

554 Figure 14: Cluster classification that was conducted with the averaged sigma SAR image and terrain
 555 products as input and six classes. The colour code is the relative accumulation rate between areas
 556 and is set as the previous classification. Compared to the previous classification, a broader and
 557 homogeneous zone of lower accumulation (orange) existed in the central valley downward from
 558 the BIA, including SP1 and SP4. The red class is restricted to specific areas close to the mountain.

559 The same flux features appeared in the zone between SP3 and SP5. We observed these
 560 features in the final classification because the TanDEM-X DEM was used. We evaluated the
 561 surface profile, and not all changes in the classes were responses to changes in the
 562 slope/aspect at the surface. We also tested terrain flattening in the image pre-processing chain
 563 by using the SNAP software with both TanDEM-X and REMA, but the additional processing
 564 did not remove these features from the SAR image.

565 4. Discussion

566 4.1. AWS and Snow Depth

567 The AWS record (2014-2017) showed no marked accumulation season, although we
568 noticed a significant episode of accumulation that began in autumn and lasted until late
569 spring. Except for 2014, most of the accumulation was deposited in very few events, similar
570 to what was observed by Reijmer & Broeke (2003) in Droning Maud Land closer to the coast
571 at lower elevation. The SMB was within the range of previous estimates. Despite the short
572 observation period, temporal variability existed, with 50% less accumulation in 2017
573 compared to 2014.

574 Previous meteorological records have been available from AWS Wx7 since 2010,
575 which was replaced by Wx14 in 2014 and is operated by the private company Antarctic
576 Logistics and Expeditions (ALE) at the Union Glacier ice runway (79°45.93' S, 83°13.58' W,
577 705 m asl). Unfortunately, no snow accumulation was recorded. Importantly, both the
578 Wx7/14 and UNION13 stations are located in the central valley, but the stations are 15 km
579 apart. The former is located in the runway in the BIA area on the northern side of the central
580 valley and upstream, where the wind is stronger. Rivera et al. (2014) used these station data to
581 describe the meteorological conditions in their work. The latter station is located close to
582 operational support at the Chilean base camp. Hence, we assumed a predominant wind
583 direction of 255° for the wind-effect index based on our records instead of 225°, as reported
584 by Rivera et al. (2014).

585 The meteorological data, despite the short collection period, provided climate trends
586 and unique in-situ data on the regional climate conditions. UNION13's location is suitable for
587 building a unique long-term record of snow accumulation with high temporal resolution. In
588 contrast to the runway's location, an area with zero or negative mass balance, UNION13 is

589 located in the central valley downward from the basin catchment and downward from the
590 BIA, where the SMB increases.

591 4.2. *Snowpit*

592 Three types of grains were present: irregular, rounded and faceted. We did not observe
593 precipitated grain types that were preserved in any stratigraphic profile, even Graupel.
594 Graupel grains are the most stable form because of their low surface/volume ratio, which
595 allows them to be preserved for extended periods in the snowpack with less metamorphism.
596 This finding indicates the dominant role of wind on depositional processes because of
597 transport or metamorphism agents. Irregular crystals are a varying crystal type that are no
598 longer recognized and are already within a grade of metamorphism in the snowpack. Rounded
599 crystals indicate a later stage of metamorphism, where the grain grows at the expense of
600 smaller particles. However, this finding also indicates drifted snow. Faceted crystals indicate
601 constructive metamorphism, which is driven by temperature gradients and water vapour.
602 Faceted crystals also indicate a layer that is exposed to wind for a longer period at the surface,
603 especially during winter, when the air humidity is low; the intense wind blowing at the
604 surface acts to remove water vapour, generating a vapour gradient that consequently favours
605 constructive metamorphism. Thus, we can interpret faceted crystals as wind-exposed
606 depositional areas.

607 This assumption agrees with the field knowledge, where we observed sites that were
608 less exposed to wind in SP2 and SP6 (Figure S2 and Figure S6), and we rarely observed
609 faceted crystals. The average small grain sizes ranged between 0.5 and 1.0 mm, and both
610 profiles exhibited fewer layers, which were usually more extensive than in other SPs,
611 indicating a higher load of deposition per event. Interestingly, we also noted a difference
612 between the two sites. Rounded forms were dominant in SP6, which is located inside the
613 tributary Driscoll Glacier, whereas irregular snow crystal types were more frequent in SP2 on

614 the opposite side of the main trunk of UG. The potential reason was the supplied mass source
615 into these two environments. In Driscoll, the snow is drifted/transported a longer distance and
616 eroded to rounded forms. In addition to a more protected area because of the U-shaped valley
617 with an orientation transverse to the predominant wind direction, we note smaller particle
618 sizes as an effect of the position at the leeward side of the mountain, as observed by
619 Ingvander et al. (2016), decreasing the wind-transport potential. The grain-size range that was
620 observed in our field data matched that in Ingvander et al. (2016) along the coastal zone in
621 Droning Maud Land. These authors found larger grain sizes compared to those in the higher
622 plateau section, where higher accumulation was observed. Additionally, the transitional-zone
623 samples and the first ascending polar-plateau samples were smaller and more homogeneous,
624 which could have caused by the position of the samples in the mountain range or on the
625 leeward side of the mountain range facing the plateau. In the same manner, our samples in
626 areas that were less exposed to wind, e.g., SP2 and SP6, exhibited smaller and more
627 homogeneous grain sizes.

628 Distinct layers of snow built up in the snowpack because of the intermittent nature of
629 precipitation, the action of wind and continuously ongoing metamorphism of snow. Each
630 stratigraphic layer differed from the adjacent layers above and below in terms of the
631 microstructure and/or density, which together define the snow type, snow hardness, and snow
632 temperature (Fierz *et al.* 2009). Therefore, we can interpret each layer as a distinct
633 depositional event at any one time because the physical and mechanical properties depend on
634 these characteristics. Palais et al. (1982) studied snow stratigraphy at Dome C and recognized
635 a thin crust as a proxy to identify annual layers because this type of thin crust is usually
636 produced at the surface in late summer and subsequently buried, although such crusts may no
637 longer be easily recognizable after firnification. We did not find evidence of the cyclic layer
638 to be well marked with a thin hard crust. At SP3 (Figure S3), two hard crusts covered a soft

639 crust (i.e., fist hardness) in the first 40 cm, each layer being 20 cm thick. However, based on
640 the amount, this finding probably represents two distinct, short events and not the majority of
641 the annual accumulation. As shown in the UNION13 AWS records, the snow accumulation
642 varied throughout the year, with sporadic depositional events and continued
643 sublimation/erosion (Figure 4). The larger number of layers in the 2-m depth interval
644 indicates several deposition events in the years with low amounts, where each layer was
645 exposed to the air-snow interface for a longer time and metamorphism more intensely
646 differentiated the layers. Thus, the annual SMB could not be estimated for every snowpit.

647 Density profiles can reveal transitions between layers through high-amplitude
648 inflections. Harper & Bradford (2003) Compared the number of layers that were identified by
649 both methods under the same sampling resolution and observed twice as many layers in a
650 density profile through a permittivity probe. The errors that were associated with sampling
651 and weighing the volume of snow were approximately 10% (Harper & Bradford 2003). In our
652 data, we observed many more stratigraphic layers than density inflections, considering the
653 low sampling resolution of 10 cm. Because of the air temperature's seasonal cycle, the
654 gradient between the snow surface and atmosphere increased by the end of summer and the
655 beginning of autumn, and densification intensified. Therefore, a high inflection of high-
656 density layers could be approximately traced as a single year's SMB, but we could not
657 confirm this point. However, we have four to five years of accumulation in the first 2 m of
658 SP2, in contrast to the five to six years of accumulation based on the age model that was used
659 by Hoffmann et al. ('in review') close to the EPCCGU. This difference can be within the
660 accuracy in the age model because of inter-annual accumulation variability. McMorrow et al.
661 (2002) highlighted the inter-annual accumulation variability at Law Dome (East Antarctica)
662 and outlined its implications for interpreting the ice-core record. The short period in the 2-m

663 snowpits does not represent the inter-annual variability, and the influence of the inter-annual
664 variability on the spatial variability cannot be isolated.

665 4.3. SAR

666 The differences that were observed in the backscattering indicated differences in
667 snowpack morphology. Rott et al. (1993) found low backscattering coefficients for areas with
668 permanent dry snow, high accumulation rates, and homogeneous snow morphology. Our
669 interpretations of SP2 and SP6 matched, and these areas displayed low backscattering in the
670 SAR images. West of the Chilean base camp EPCCGU was a brighter patch that appeared as
671 a wind track (red arrow in Figure 7), which originates from the small tributary to the
672 southwest of the "Criosfera Glacier". Consequently, the higher backscattering suggests a zone
673 of higher density and grain size. The interpretation of a wind track makes sense, whereas a
674 darker patch of a wind-protected zone was present on the eastern side, which was elected by
675 ALE and Chile as a base-camp site.

676 We investigated the enhanced glacial-flow structure in the SAR images. Some of the
677 contrasting areas that changed from low to high backscattering were followed by changes in
678 the surface elevation in the TanDEM-X. We initially believed that the explanation could be
679 the alignment of the surface aspect at a right angle to the SAR antenna (the satellite azimuth
680 was from right to left and nearly parallel to the valley flow). The surface aspect reduced the
681 incident angle, increasing the surface-scatter contribution. For frequencies below 10 GHz,
682 scattering losses are neglected, and volume scatters are dominant (Rott et al. 1993). The main
683 argument resulted from the low contrast between layers (Du *et al.* 2010). Forster et al. (1999)
684 quantified a volume scatters contribution of 100% (>95% for 25°) with an incident angle
685 greater than 30°. We checked the incident angle in the geometrically corrected SAR image,
686 and the angle varied only in the range of 22-26° between areas of low and high backscattering.
687 As we presented in the results based on the SAR simulated image, the surface relief was

688 smoother than what was represented by TanDEM-X. Therefore, TanDEM-X were
689 representing some changes in the subsurface. We examined a GPR transect that crossed these
690 zones and found that the backscattering varied with the snow-ice horizon depth (data not
691 published yet). The backscattering increased where the snowpack was thinner than 10 m deep
692 because some of the signal is reflected by the ice back to the snowpack, increasing the volume
693 scatter. The terrain-flattening process, which normalizes the returning signal by using locally
694 illuminated areas, did not smooth these features in the SAR images.

695 Another fact that could have enhanced the flow features would be converging fluxes
696 from tributaries enhancing the contrast vertically and horizontally between deeper layers,
697 increasing the multilayer scattering. According to Tsang et al. (2006), multiple scattering can
698 raise signals by a few decibels if the scattering albedo is close to one. Dierking et al. (2012)
699 found few test sites with albedo values larger than 0.7 in the C-band and 0.8 in the Ku-band.
700 In these cases, the accumulation rates were low, and high scattering albedo was located at
701 greater depths (caused by larger grains). At specific locations in Greenland and Antarctica,
702 deep hoars formed at the onset of summer (comparatively larger grain sizes, 2-5 mm). In
703 addition, the northern side of the valley had a shallower firn layer, and we expect a higher
704 density gradient in the first ten meters within the SAR signal.

705 Comparing the snow density and grain-size maps with the field data did not indicate
706 absolute correspondence. The applied algorithm overestimated the density over the snowpit
707 density (Figure 10). The field measurements corresponded to the first 2 m and were averaged
708 from the 10-cm measurements without considering the thickness of each layer. The
709 attenuation depth of the X-band can reach 10 m in areas of dry snow (Rott *et al.* 1993, Wessel
710 *et al.* 2016), so the signal probably interacts with denser layers beyond the first 2 m, although
711 Espinoza et al. (2014) estimated that most backscattering (<95%) occurs in the initial 2 m of a
712 package. A density profile that was derived from a firn core (indicated by a green dot in

713 Figure 10) close to SP2 showed densities from 400 kg m^{-3} at the surface to 550 kg m^{-3} at a
714 depth of 8 m (Hoffmann et al. 'in review'), which was more similar to the algorithm-derived
715 density than the average field density. The density map alone did not appear to explain the
716 variations in the snowpack; for example, SP3 and SP5 are interpreted as being in high-density
717 (SAR-derived) areas because of high backscattering, but other characteristics, such as the
718 grain size and number of layers, likely contribute to the high backscattering in addition to the
719 density. For example, SP3 and SP5 had larger grains and more layers than SP2 and SP6,
720 indicating lower accumulation rates than those in SP2 and SP6. A lower accumulation rate
721 will tend to develop larger grain sizes for snow because each snow layer is exposed at the
722 surface for a longer period (Linow *et al.* 2012).

723 In addition, the map indicated that SP2 and SP6 had the same density values of
724 approximately 440 kg m^{-3} , but the field densities were 417 kg m^{-3} and 338 kg m^{-3} ,
725 respectively. Similarly, SP1 and SP4 also had density values in the same range of $520\text{-}570 \text{ kg}$
726 m^{-3} , but the field density differed from 433 kg m^{-3} to 400 kg m^{-3} , respectively. Despite this
727 disagreement between the two datasets, the mean density of the 2-m snowpits matched the
728 stratigraphic interpretation if we considered the southern side of the valley (SP1, SP2 and
729 SPA) separately from the northern side (SP3, SP4, SP5, and SP6). On the southern side, SP1
730 and SP2 differed in terms of density, corresponding to a higher density at SP1 because of
731 wind compaction and a lower density at SP2. On the northern side, the density range differed
732 from that on the southern side, but a gradient existed, with higher values in SP4 decreasing to
733 those in SP3, SP5 and SP6. These differences were likely influenced by other characteristics,
734 such as the grain type and wind transport between the southern and northern sides.

735 SPA had a low mean density, and many layers could explain the higher backscattering.
736 We hypothesized that SPA is located at a wind-protected site (white circle in Figure 10)
737 leeward from Rossmann Mountain, where the amount of snow that accumulates originates

738 from blowing snow that bypasses the topographic barrier, which reflects the small grain size
739 that was observed in the stratigraphic analysis. Because the amount of snow in each
740 depositional event was small, the snow was metamorphosed in the first centimetres because of
741 the longer exposure time, creating a layered snowpack. On the northern side of the central
742 valley, SP4 exhibited approximately the same backscattering value as SP1 but a lower
743 density, which was caused by hypothetical wind compaction that was lower than that on the
744 other side of the valley. Compared to SP4, SP1 showed more hard layers in the first meters.
745 SP4 also showed hard layers, which were intercalated with softer layers. The lower density at
746 SP3 downwind from SP4 could indicate the drifting of snow from SP4 to SP3, increasing
747 accumulation at SP3. The greater number of layers compared to SP4 and the large grain size
748 could explain the high level of backscattering at this site.

749 *4.4. Terrain products*

750 The aspect map clearly contrasted the right side of the wind track with intercalating
751 bright values (1), indicating pixels with windward orientations (Figure 12c), which extended
752 to SP2. However, the darker colours (-1) along the wind track indicated a leeward aspect. This
753 difference explained the change in accumulation between SP1 and SP2: leeward snow tended
754 to be carried away, whereas windward surfaces tended to accumulate snow. Goodwin (1990)
755 showed that accumulation rates are higher on windward slopes than leeward slopes. In fact,
756 SP2 and SP6 were located in areas with a windward aspect, while SP1 had a leeward aspect.
757 However, SP4 contrasted what was expected in terms of the windward aspect, and SP5 had a
758 leeward aspect. These slope values explain why SP4 (SP5) had lower (higher) accumulation
759 even in a windward (leeward) area. The slope values for both locations were lower than those
760 for the other sites (Figure 12d).

761 Accumulated snow is deposited in the form of surface microrelief as topographic
762 features on spatial scales of 10-100 m (Goodwin 1990). These features reflect higher slope

763 values, corresponding to areas where we expect higher accumulation. We observed a low-
764 slope area that extended from SP1 in the wind-track zone, crossing the valley to the
765 surrounding area of SP4 (Figure 12d.). We interpreted this zone as a flat surface because of
766 wind action with lower accumulation. The roughness was similar to the slope. Higher surface
767 roughness favours snow deposition in intercalating high- and low-pressure surfaces,
768 promoting turbulent air fluxes (Frezzotti *et al.* 2002).

769 The wind-effect algorithm considers a fixed mean wind direction for the entire grid
770 and might not perfectly represent local conditions that are related to katabatic winds, which
771 tend to follow the topography. The algorithm also depends on the search distance; varying this
772 distance enables us to obtain more or less wind shelter in the tributary of the Schanz and
773 Driscoll glaciers. In the Driscoll valley, as observed at SP6, the snow deposition was
774 dominated by small grains that bypassed hills and were further deposited. Using a high search
775 distance in the wind-effect algorithm ignores the adaptability of wind streams to the land
776 surface (Böhner and Antonic 2009). The 255° wind-effect map corresponded better to the
777 main trunk valley direction and therefore represents the effect of katabatic wind moving down
778 the glacier (Figure 12b). This map adequately represents intercalated zones that are more and
779 less exposed to wind down the valley. The wind-effect field was also more exposed at SP4.
780 SP5 and SP3 were located directly adjacent to an area that was exposed to wind. These
781 locations mean that these areas received blown snow from wind-exposed areas. Figure 12b
782 does not show a higher wind exposure at SP1 or a lower exposure at SP2, as we would expect.
783 However, we interpreted the algorithm results as generally exposed to the wind, which means
784 that the algorithm tended to interpret the surface that was sloped windward as exposed and the
785 surface that was sloped leeward as sheltered. This determination makes sense, for example, at
786 the border outside the masked area with low backscattering values in the SAR image (Figure
787 7). This area is located directly beyond a wind-exposed area in the BIA and receive blowing

788 snow, similar to SP5 and SP3. On a local scale, the algorithm could not consider the effect of
789 the slope and topography when modelling wind-direction changes because the algorithm
790 considers only a fixed wind direction. Most likely, local features such as the wind that flow
791 down through the smaller valleys and throats were not modelled. For example, we note the
792 slight difference between the mean wind directions that were reported by Rivera et al. (2014)
793 at the runway and in our work at the UNION13 AWS.

794 4.5. Cluster

795 We masked 32% of the 1687-km² imaging area as mountainous/sloped area and BIA,
796 both of which showed low and negative accumulations, respectively. Approximately 41% of
797 the area was classified as high accumulation and 28% was classified as low accumulation
798 (Table IV). Most of the ablation stakes that were used to infer the mass balance of Union
799 Glacier in previous studies (Rivera et al. 2014) were located in areas that were considered to
800 be low accumulation, which would suggest that the mass balance could be higher than
801 previously thought if we considered zones with high accumulation rates. Field observations
802 indicated that this difference could reach 0.1 m w.e. a⁻¹, affecting the net mass balance for the
803 area by as much as 0.041 m w.e. a⁻¹. Further work will investigate a 72-km-long profile along
804 the glacier, which will provide quantification and comparisons between the different
805 depositional zones. The results help to identified different area and can guide future works on
806 the attenuation depth in dry snow to better correct new DEMs that are derived from TanDEM-
807 X interferometry (Wessel *et al.* 2016).

808 Table IV: Area of each cluster class and percentage of the total imaging area.

	km ²	%
Masked	536.4	31.8
+++ Accum.	175	10.4
++	254.5	15.1
+	255.6	15.1
-	151.2	9
--	255.9	15.2
--- Accum.	59	3.5
TOTAL	1687.4	100

809

810 If we consider this dynamic for the entire Ellsworth Mountain Range, which comprises
811 four significant basins from MEASURE Antarctica Boundaries, including the Rutford,
812 Minnesota, Union, and Hercules glaciers, we have a total area of 85,800 km². Considering the
813 central glacial valleys exclusively as the study area, we found that ~12,000 km² (6,000
814 Sentinel Range and ~6,000 Heritage Range) of the total 85,800-km² area should have the
815 same high spatial variability of snow accumulation. Thus, we would have a masked area of
816 roughly 2,048 km² from the ADD rocky area (or only a 378-km² base on the Landsat 8 rocky
817 area), 35 km² from the ADD moraine area and 612 km² of BIA, which corresponds to 22%.
818 An example of the importance of considering the variability of SMB was demonstrated by
819 (Frezzotti *et al.* 2004) in East Antarctica. High spatial variability was observed because of
820 wind-driven sublimation; consequently, previous SMB maps that did not consider these
821 factors overestimated the SMB.

822 **5. Conclusions**

823 Our results showed that wind-exposed areas had larger snow grains (i.e., 1-4 mm
824 versus 0.5-1 mm in wind-protected areas), faceted forms from greater exposure to the
825 temperature gradient, more deposition layers and layers with greater hardness. The
826 densification processes in these areas were more intense and produced thicker hard layers.

827 The stratigraphic profiles and the density and grain-size maps confirmed a distinct pattern of
828 snowpack characteristics, which further indicated particular depositional rates along the
829 glacier. The field interpretations guided us to delimit different depositional zones, and an
830 initial cluster analysis successfully classified SAR backscattering into six classes. Some zones
831 were classified as low accumulation-rate zones because of high backscattering, although the
832 field data suggested a higher accumulation rate. We introduced terrain products to better
833 isolate these zones in the cluster classification, which were directly related to or influenced
834 the depositional process. Roughness data were a good indicator of the deposition dynamics.
835 The wind effect was limited to representing local wind flow and provided only a general
836 scenario but failed to explain the local variations in the depositional zone. In addition to wind
837 exposure, we had to consider where the wind was blowing. If the wind was channelled
838 through throats without lifting snow for later deposition, the wind would erode and flatten the
839 surface into a wind-glazed surface, as observed in the wind-track area close to SP1.

840 The SAR image significantly enhanced the glacier-flow features, especially in the zone
841 of converging fluxes. A simulated SAR image was generated with the TanDEM-X DEM and
842 the REMA DEM. The later did not produce such features, indicating that the cause of the high
843 backscattering may have been in the subsurface. The features probably reduced the volume-
844 scatter contribution, increasing surface reflectance in the multi-layered media; thus, the SAR
845 signals in such zones must be carefully interpreted.

846 The imaged area focused on the mountainous area of the glacial basin, and up to 40%
847 of the areas were masked. This high percentage indicates that a significant area might not
848 represent the mean SMB from the coarse-resolution data. Moreover, these masked areas
849 corresponded to 172 km² of BIAs (10%), where we expected to find a negative mass balance.
850 The other component was mountainous/rocky areas with a steep slope and likely lower rates

851 of accumulation. These areas represented 469 km² (16%) of the total basin area of 2955 km²,
852 or 690 km² if we considered the MEASURE Antarctica Boundaries.

853 These results should assist future investigations of SMB variability and could
854 influence or act as a significant factor in interpolating this variable in climate models. These
855 results can also guide future works on the attenuation depth in dry snow to better correct new
856 DEMs that are derived from TanDEM-X interferometry. Future work will investigate the
857 accumulation rate by using GPR. The accumulation rate could be calibrated by comparing
858 AWS snow-depth data with annual snowpit and GPR profiles. A total of 72 km of GPR
859 profiles will facilitate the quantification of the accumulation variations along the glacier.

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867 elevation models, respectively.

868 **Description of the authors' responsibilities**

869 JAN and RJ collected the field data. CFG processed the data and wrote the major
870 components of the paper. CFG and JAN jointly led the study. All the authors contributed to
871 and revised the manuscript at all stages.

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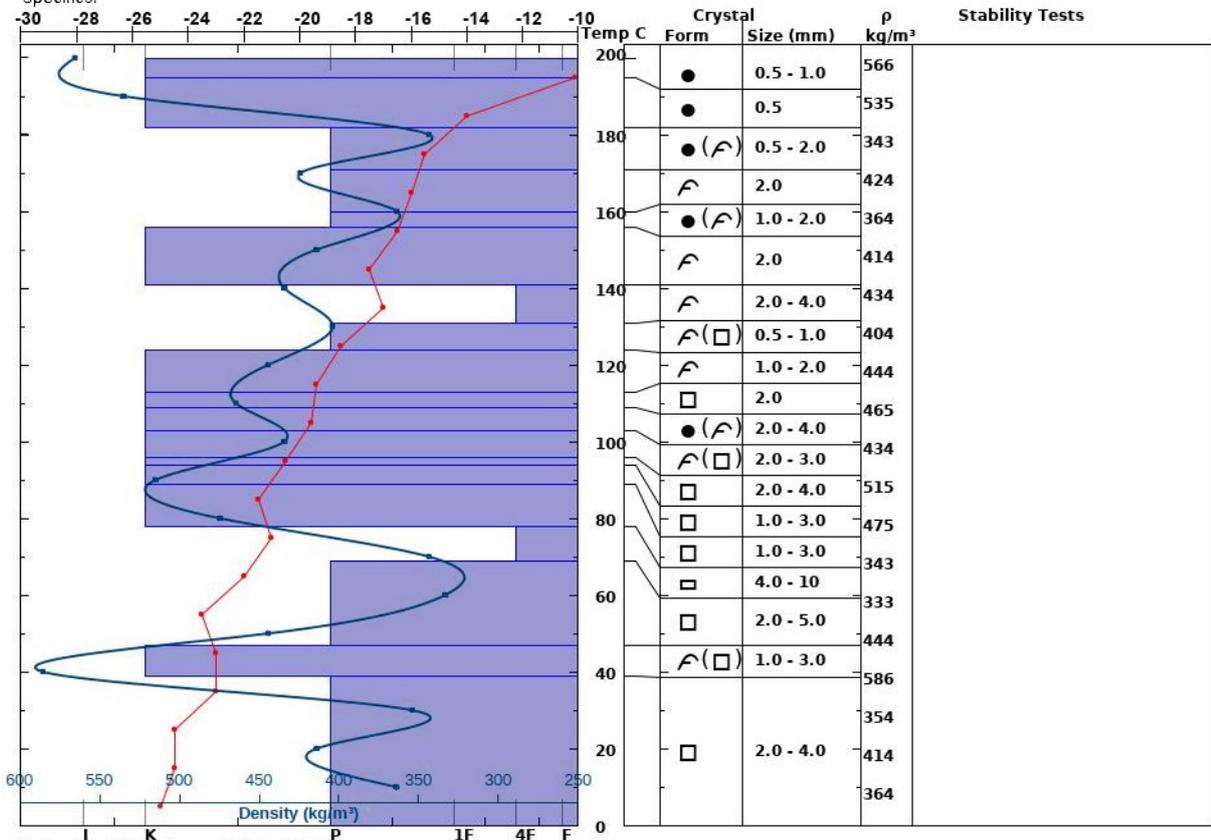
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SUPPLEMENTARY MATERIAL

Snow-deposition characteristics from SAR and geospatial analysis at Union Glacier, Antarctica

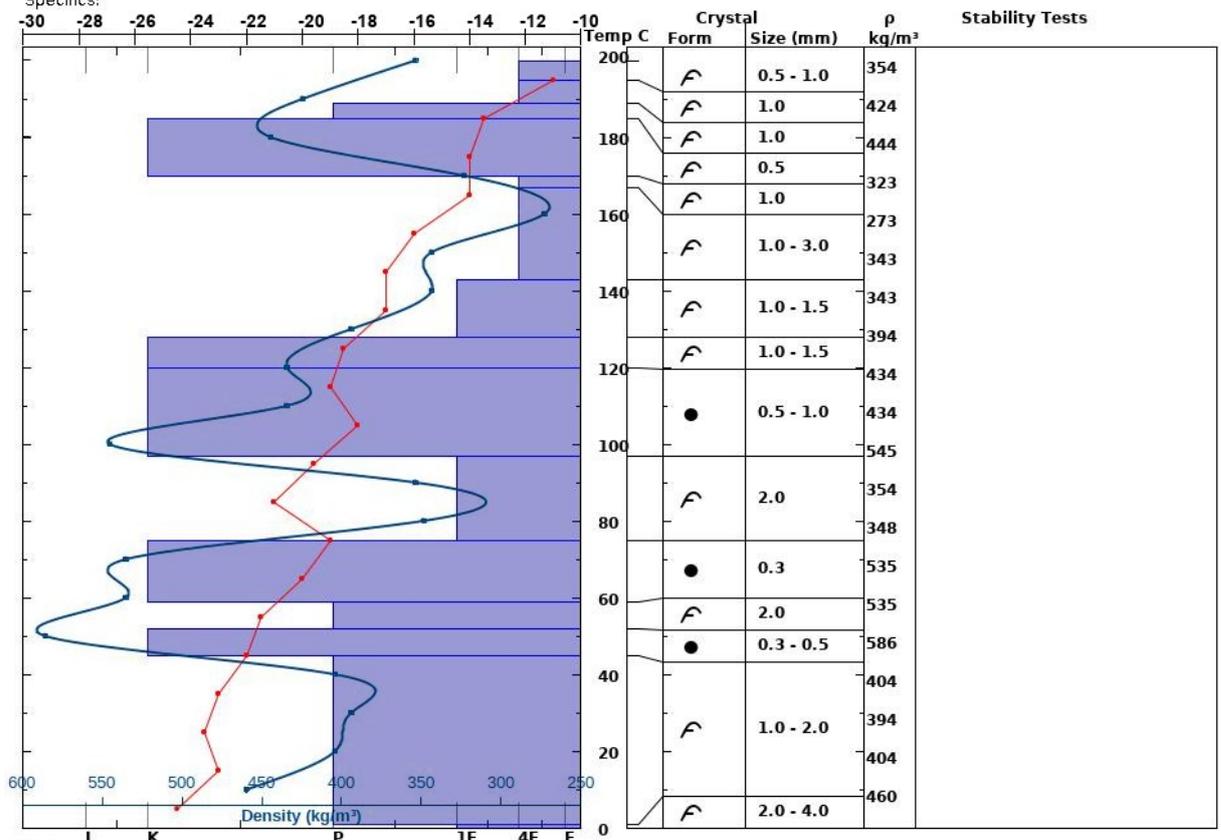
Christian Florian Göbel (cfgobel@gmail.com), Jorge Arigony-Neto, Ricardo Jaña, Rodrigo Gomez Fell, Jean de Almeida Espinoza, Francisco Fernandoy, Ian D. Goodwin, Gulab Singh

Snow Pit Profile Observer: **Christian Gobel** Stability on similar slopes: Stability Test Notes: Layer notes:
SP1 **Sat Dec 31 14:25:00 BRST 2011** Air Temperature: **8 C** **69-78: capped Column**
Ellsworth, Antarctic Co-ord: **79.79227 S 82.968055 W** Sky Cover:
Elevation (m): **718** Slope:
Aspect: Wind loading: Precipitation:
Specifics: Wind:



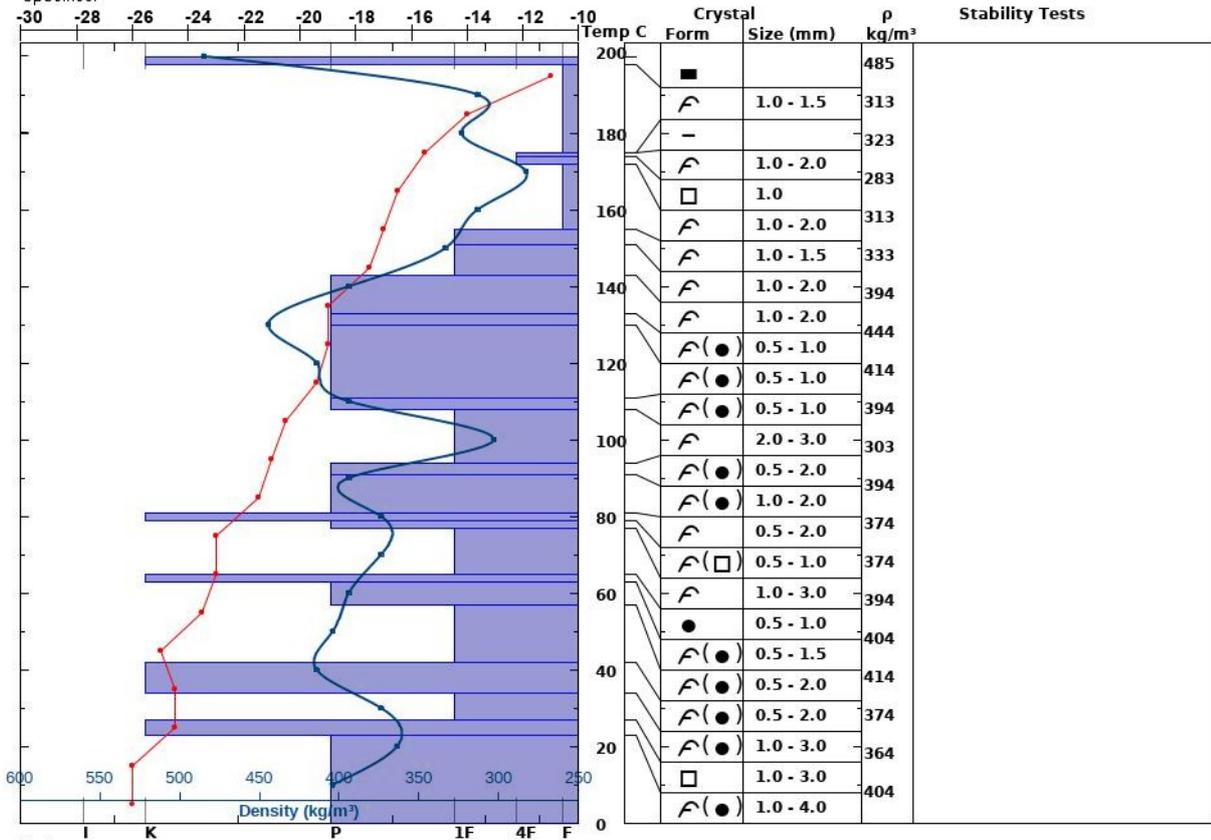
Supplementary Figure S1: Snowpit 1 diagram. Temperature profile plotted in red with the top axis. The bars indicate hand hardness with the bottom axis. Density profile plotted in blue with the axis at bottom above the hand hardness axis.

Snow Pit Profile Observer: **Christian Gobel** Stability on similar slopes: Stability Test Notes: Layer notes:
SP2 **Thu Dec 29 15:00:00 BRST 2011** Air Temperature: **5.9 C**
Ellsworth, Antarctic Co-ord: **79.79227 S 82.968055 W** Sky Cover:
Elevation (m): **598** Slope:
Aspect:
Wind loading:
Precipitation:
Wind: **Calm**



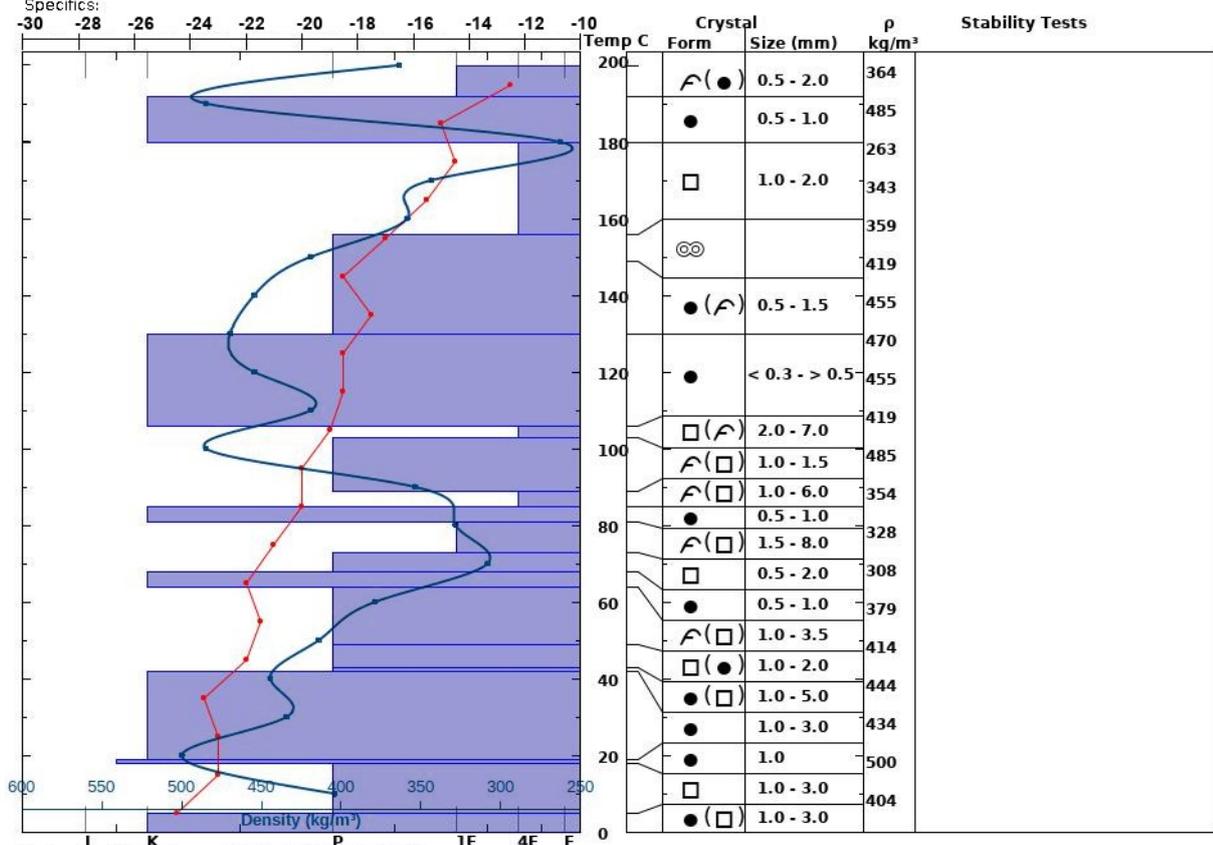
Supplementary Figure S2: Snowpit 2 diagram. Temperature profile plotted in red with the top axis. The bars indicate hand hardness with the bottom axis. Density profile plotted in blue with the axis at bottom above the hand hardness axis.

Snow Pit Profile Observer: **Christian Gobel** Stability on similar slopes: Stability Test Notes: Layer notes:
SP3 **Sat Apr 02 13:07:00 BRT 2011** Air Temperature: **8.8 C**
Ellsworth, Antarctic Co-ord: **79.79227 S 82.968055 W** Sky Cover:
Elevation (m): **554** Slope:
Aspect: Wind loading: Precipitation:
Specifics: Wind:

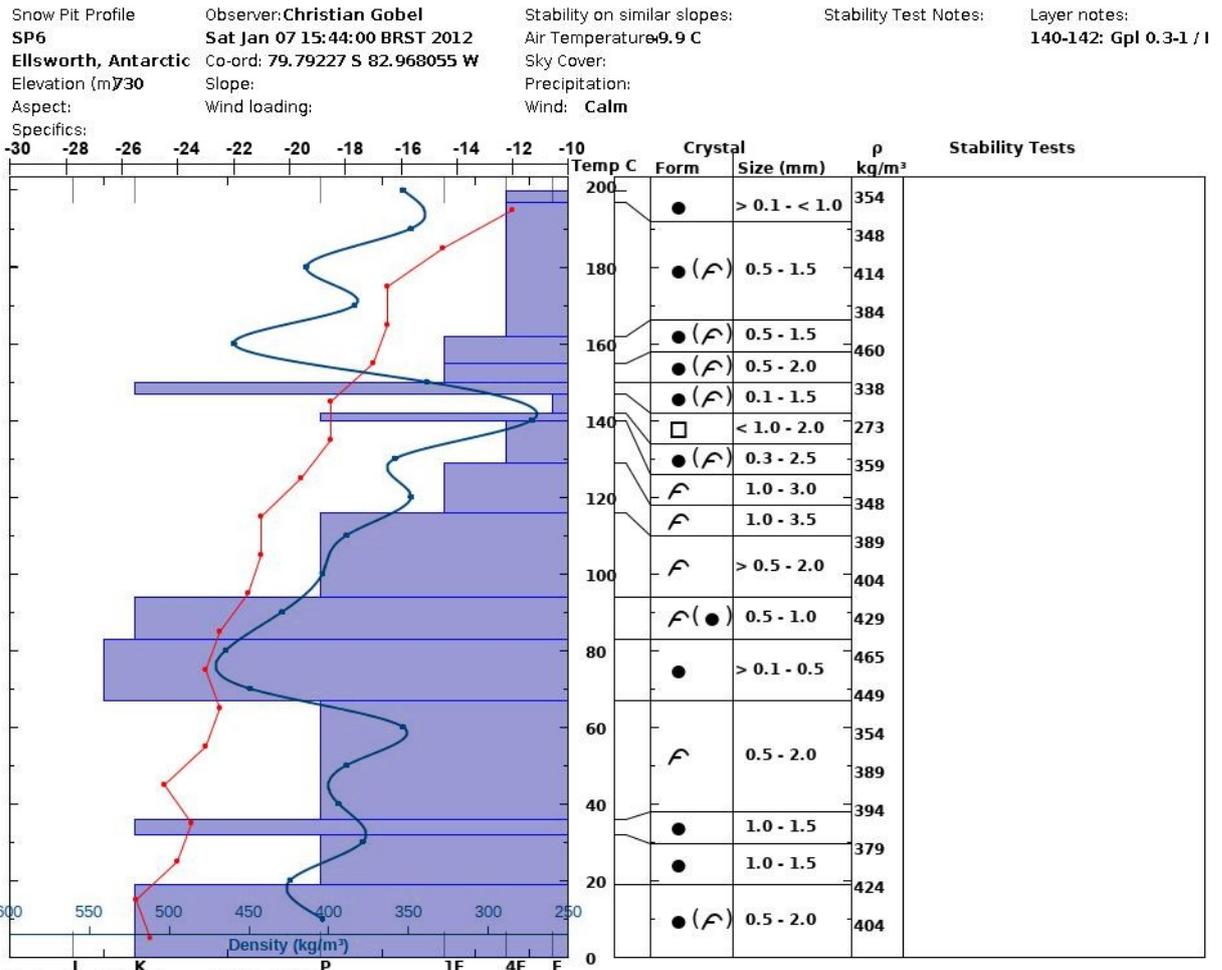


Supplementary Figure S3: Snowpit 3 diagram. Temperature profile plotted in red with the top axis. The bars indicate hand hardness with the bottom axis. Density profile plotted in blue with the axis at bottom above the hand hardness axis.

Snow Pit Profile Observer: **Christian Gobel** Stability on similar slopes: Stability Test Notes: Layer notes:
SP5 **Fri Jan 06 14:14:00 BRST 2012** Air Temperature: **7.5 C** **149-149: Linha de d**
Ellsworth, Antarctic Co-ord: **79.79227 S 82.968055 W** Sky Cover: **73-81: cristais gigar**
Elevation (m): **569** Slope: Precipitation:
Aspect: Wind loading: Wind: **Light Breeze**

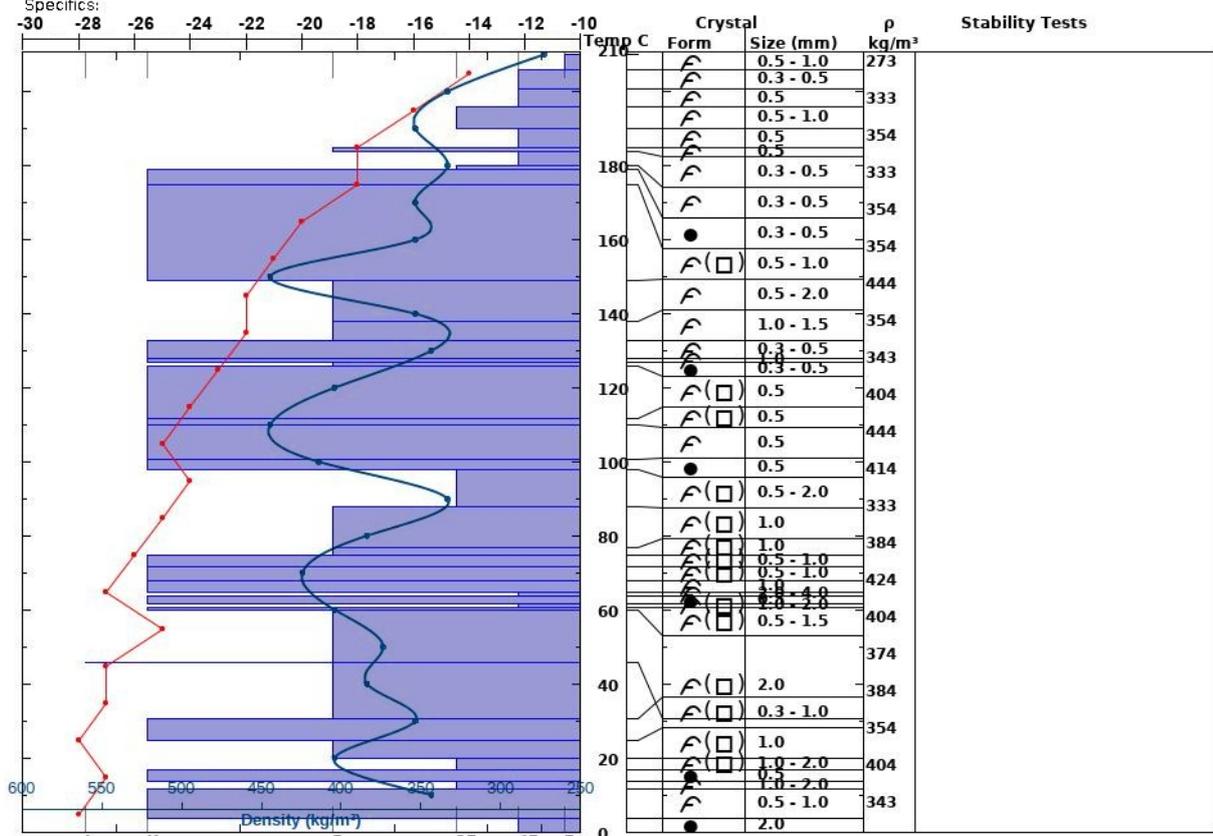


Supplementary Figure S5: Snowpit 5 diagram. Temperature profile plotted in red with the top axis. The bars indicate hand hardness with the bottom axis. Density profile plotted in blue with the axis at bottom above the hand hardness axis.



Supplementary Figure S6: Snowpit 6 diagram. Temperature profile plotted in red with the top axis. The bars indicate hand hardness with the bottom axis. Density profile plotted in blue with the axis at bottom above the hand hardness axis.

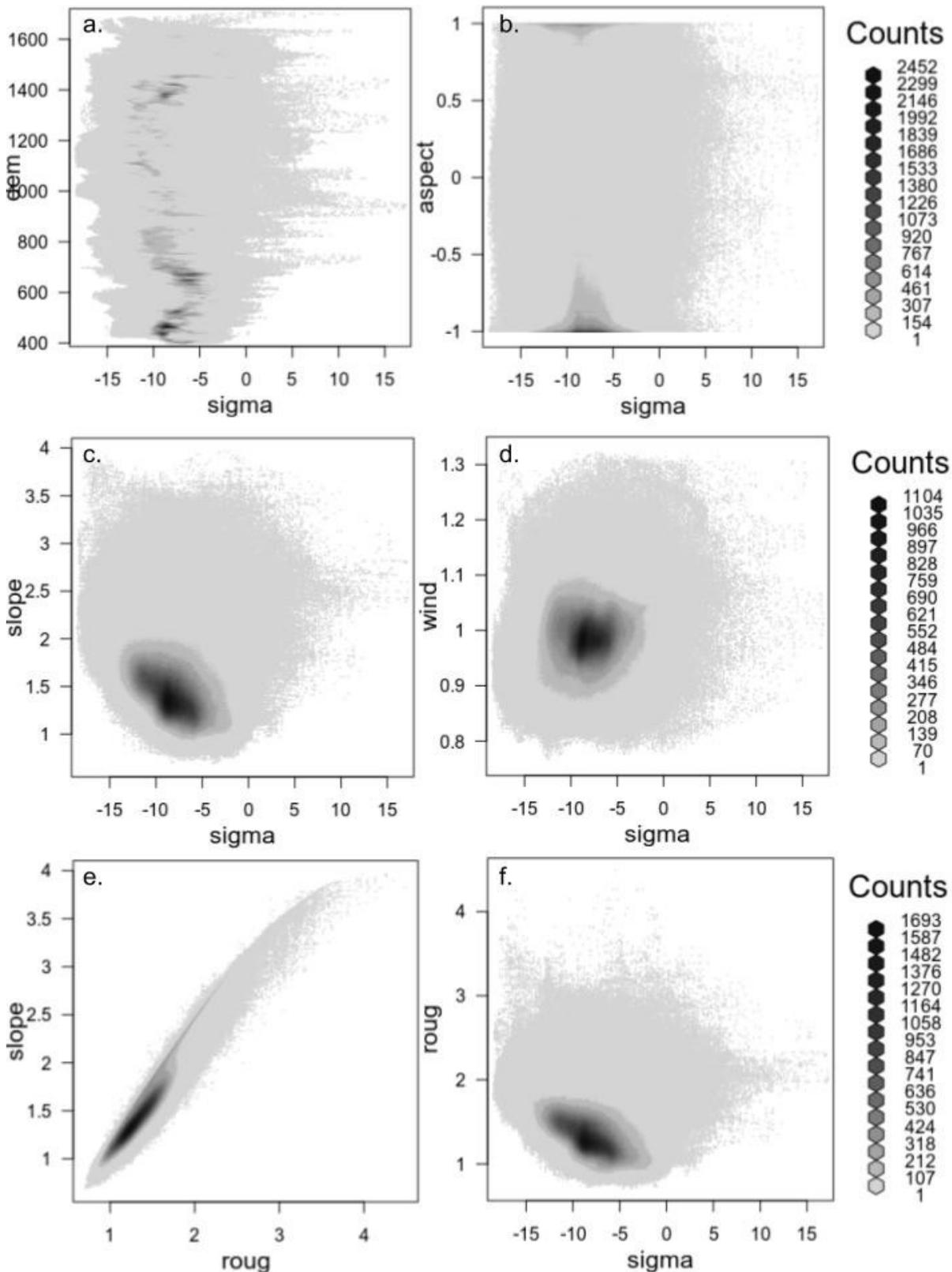
Snow Pit Profile Observer: **Christian Gobel** Stability on similar slopes: Stability Test Notes: Layer notes:
SPA **Mon Dec 19 18:43:00 BRST 2011** Air Temperature: **10.4 C** **184-185: Crosta**
Ellsworth, Antarctic Co-ord: **79.79227 S 82.968055 W** Sky Cover: **179-180: Crosta**
Elevation (m): **720** Slope: Precipitation: **46-46: Hard line**
Aspect: Wind loading: Wind: **Light Breeze**



Supplementary Figure S7: Snowpit A diagram. Temperature profile plotted in red with the top axis. The bars indicate hand hardness with the bottom axis. Density profile plotted in blue with the axis at bottom above the hand hardness axis.

Scatter Plot - Hexbins

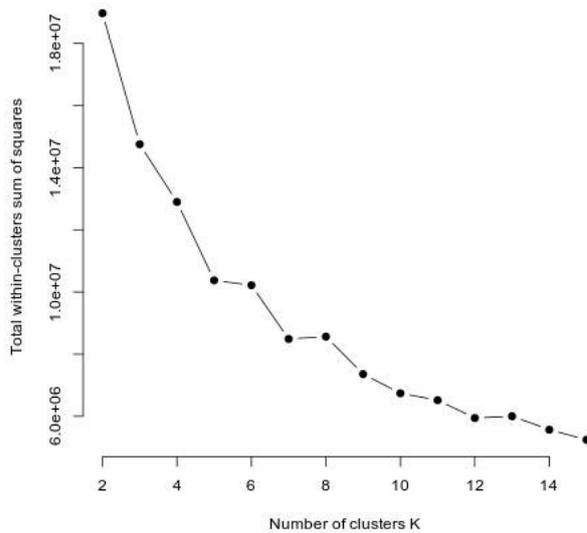
We plotted each variable against sigma in hexbin scatterplot because of the high density of data. No clear cluster of values is observed.



Supplementary Figure S8: Scatterplot in hexbins for each terrain product against sigma values, except for (e.) which is slope against roughness. (a.) elevation values; (b.) relative surface aspect to azimuth of the prevailing wind direction where, 1 correspond windward and -1 leeward aspect; (c.) slope expressed in cubic square root; (d.) wind effect where, values <1 indicate wind sheltered and >1 wind exposed areas; and (f.) roughness also expressed in cubic square root.

Elbow test

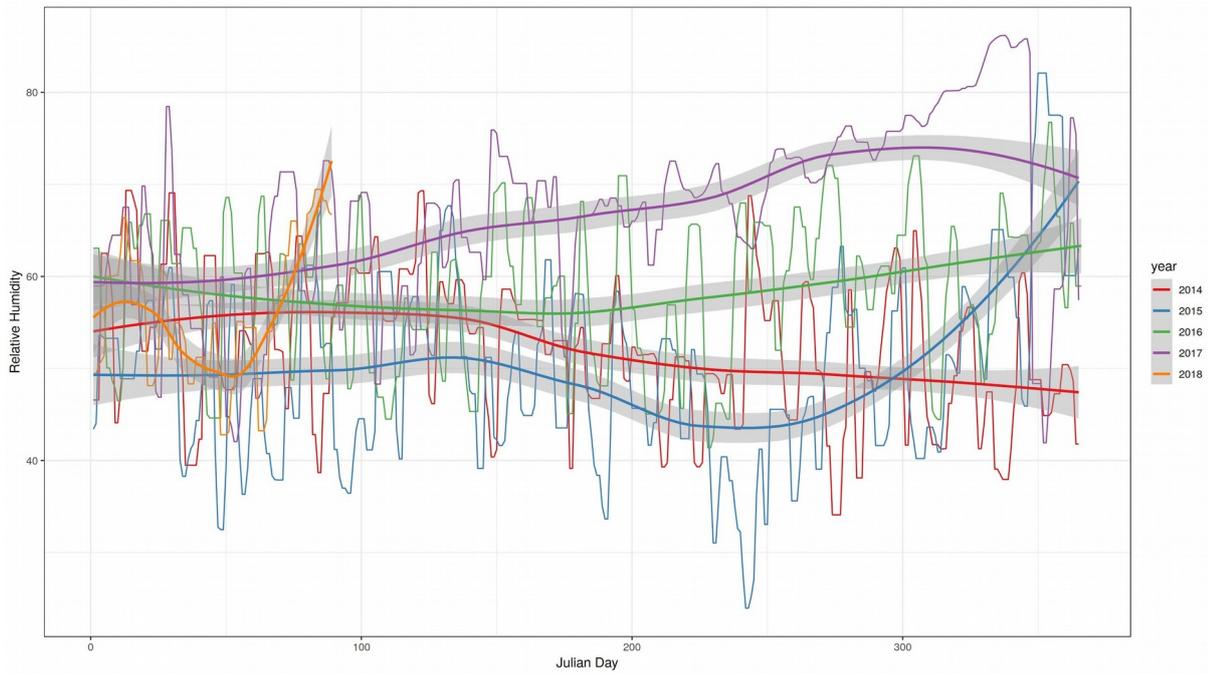
We did the Elbow test selecting only roughness, slope, dem and sigma. The result do not present a clear elbow:



Supplementary Figure S9: Plot of the sum squared error (SSE on the y-axis) for all the pixels in the clustering analysis with roughness, slope, dem and sigma as input and the number of cluster groups k (x-axis). The elbow method suggests that the number k should be at the point where increasing k provides little return when decreasing the SSE.

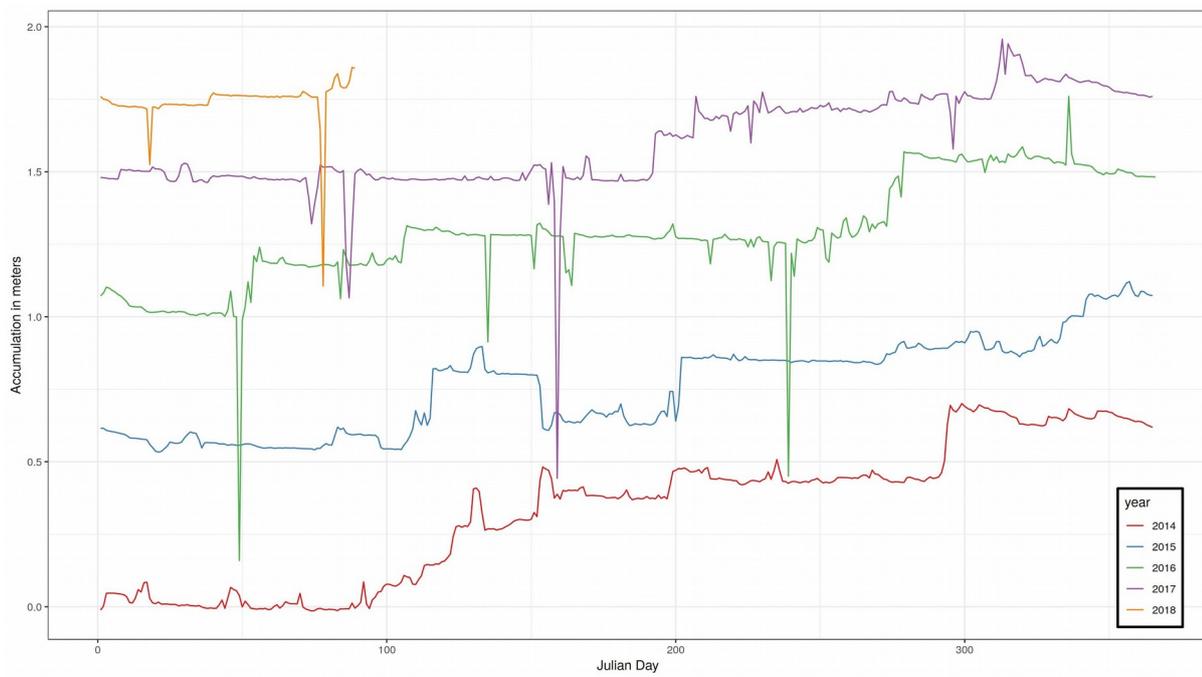
Relative humidity

From UNION13's AWS



Snow depth from SR50 sensor

Data from UNION13's AWS.



4 Capítulo 4 - Artigo 2

O segundo manuscrito, de autoria de Christian Florian Gobel, Juliana Costi, Ricardo Jaña e Jorge Arigony-Neto, é intitulado “**Influence of snowpack characteristics on TanDEM-X DEM - validation with REMA and field datasets acquired on the Ellsworth Mountains, Antarctica**”, submetido na revista *Geophysical Research Letters*

Neste segundo artigo utilizou-se dois modelos digitais de elevação distintos, comparando-os, para identificar zonas de distintas características do pacote de neve, principalmente no que se refere à taxa de acumulação. Além disso, validou-se os DEMs com dados RTK de campo. A comparação com o segundo modelo baseado em SAR, identificou zonas em que o sinal apresenta uma penetração e interação mais rasa ou profunda com o pacote de neve. Sendo assim, propõe-se o uso dos dois modelos como identificadores da variabilidade espacial na taxa de acumulação da neve.

2 **Influence of snowpack characteristics on TanDEM-X DEM - validation**
3 **with REMA and field datasets acquired on the Ellsworth Mountains,**
4 **Antarctica**

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15

16 **Key Points:**

- 17 • REMA validation with field GPS shows that elevation is overestimated by an average of
18 one meter over Union Glacier, Ellsworth Mountains.
- 19 • TDX presents higher errors. The deviation is related to snow/ice cover and could be used
20 as a proxy for snow accumulation assessment.
- 21 • In Blue Ice Areas, the offset is positive and could indicate a negative SMB. Glacial ice up
22 to 6 m deep influences the SAR backscattering.

23

24 **Keywords:** SAR; GPR; snow accumulation; glacier surface mass balance; blue ice area.

25 **Abstract**

26 The penetration and interaction of X-band synthetic aperture radar (SAR) with snowpack
27 depends on the snow layers physical characteristics related to snow accumulation processes. We
28 use the new Reference Elevation Model of Antarctica (REMA) as a reference surface to subtract
29 from the TanDEM-X elevation model (TDX) and evaluate the X-band interferometric bias in dry
30 snowpack. We confirm the REMA's high accuracy with 70-km-long geodetic measurements on
31 Union Glacier in the Ellsworth Mountains. A mean error of 1.01 ± 0.61 meters was found. TDX
32 presented a higher mean error of 2.05 ± 2.37 m. We demonstrate that the TDX surface covaries
33 with ice depth and accumulation layering changes in the GPR profiles. Furthermore, we propose
34 that both DEMs' data can be used to investigate the subsurface feature changes and ultimately,
35 the accumulation dynamic changes. Negative (positive) differences indicate high (low or
36 negative) accumulation rate areas where deeper (shallower) penetration occurs.

37

38 **Plain Language Summary**

39 A key component of understanding the mass balance of Antarctica is surface snow
40 accumulation because of high continental areas and spatial variability. Field measurements are
41 challenging to obtain, and any remote approach aids in understanding this process. We compare
42 the last two high-resolution elevation models available for Antarctica: REMA and TandDEM-X.
43 The first model is derived from visible range satellite images and therefore represents the surface
44 elevation. The latter model is based on radar interferometry at the X-band wavelength, which
45 penetrates and interacts with dry snow. The accumulation rate influences how deep the signal
46 penetrates and consequently offsets the elevation value of the elevation model. We propose a
47 comparison between both data types as a proxy to track areas with distinct accumulation
48 dynamics. Negative (positive) differences indicate high (low or negative) accumulation rate areas
49 where deeper (shallower) penetration occurs.

50

51 **1 Introduction**

52 A sufficient elevation model for Antarctica is highly necessary to improve the
53 understanding of the physical processes that influence glacier dynamics. On a large scale, surface
54 elevation affects the flow of glaciers and is crucial for atmospheric circulation modeling. The
55 resolutions of continental digital elevation models (DEM) for Antarctica have been improved in
56 the last decades from hundreds of meters, e.g., RADARSAT Antarctic Mapping Project (RAMP)
57 DEM, to tens of meters from stereoscopic medium resolution sensors, e.g., ASTER and ALOS
58 sensor derived global DEM. However, there remains a lack of data and sufficient accuracy for
59 regions inside the continent where homogeneous areas with low feature contrasts occur.

60 A very high-resolution DEM can represent surface roughness, which influences the
61 deposition and redistribution of snow. The topographic relief can also indicate the depositional
62 characteristics of the surface and subsurface (M. Frezzotti et al., 2002; Goodwin, 1990). Wind

63 plays an important role in the spatial distribution of snow and determines the surface roughness,
64 which is a consequence of the accumulation pattern (Massimo Frezzotti et al., 2004, 2007).

65 Synthetic aperture radar (SAR) interferometry-derived DEM presents both high resolution
66 and accuracy. A drawback of the method is the signal penetration on the dry snow surface and the
67 need to precisely correct the depth of the interferometric center point. The penetration of the SAR
68 signal in the snow depends mostly on the dielectric constant, in addition the band frequency,
69 which varies according to density, grain size and layering (Forster et al., 1999; Rott et al., 1993;
70 Tsang et al., 2006). These characteristics are directly linked to the accumulation rates in a
71 specific zone (Dierking et al., 2012), and in complex topographic areas such as the Ellsworth
72 Mountains, high spatial variability occurs. Thereafter, a single constant correctness value cannot
73 be expected for a broad area. Wessel et al. (2016) evaluated the TanDEM-X elevation model
74 (TDX) for Greenland with ICESat and found that the SAR penetration was up to 10 m, and the
75 resulting DEM should represent the X-band reflective surface. In other words, the resulting
76 interferometric SAR measurements represent the surface elevation corresponding to the mean
77 phase center of the backscattered signal (Rizzoli, Martone, Gonzalez, et al., 2017). A constant
78 bias is calculated and applied based on the mean height difference between the TDX and ICESat
79 elevations within selected fixed boxes in areas of homogeneous backscattering. Beginning from
80 these fixed blocks, all other Antarctica acquisitions are adjusted by relying solely on tie points
81 and previously calibrated areas.

82 The recently released REMA DEM is derived from very high spectral resolution images
83 by stereoscopy, and therefore, these data trustworthy representations of the surface-independent
84 surface cover characteristics. We test our first assumption by validating the DEM with field
85 geodesic measurements. We compared both DEMs to identify the zones where the TDX
86 presented penetration depths that are deeper or shallower than the constant bias parameter applied
87 in the TDX product generation. We proposed that the TDX-REMA approach can retrieve the
88 spatial variability of snow accumulation rates, allowing us to delimit and quantify zones with
89 higher accumulation rates, especially where the SAR signal penetrates deeper in the snowpack.

90 **2 Materials and Methods**

91 *GPS data.* Geodetic measurements were taken during the 2014 summer campaign with
92 the kinematic Global Positioning System (KGPS) method using Leica® equipment. The global
93 positioning system (GPS) points were postprocessed using Precise Point Positioning (PPP) from
94 the base station installed at the EPCCGU base camp. The track points were collected during field
95 displacement on a snowmobile with velocities ranging from 10-20 km h⁻¹. We excluded all points
96 with planimetric, altimetric or absolute accuracies greater than one meter. We snapped the dense
97 point dataset to the nearest neighboring pixel of REMA's common grid. Pixels with varying
98 numbers of points were simplified to a single mean value. The majority of pixels had only two
99 GPS points fused, and 85% had only 3 points.

100 *DEM accuracy assessment.* We used the GPS data to validate the REMA accuracy, as
 101 we wanted to use it as a reference surface for comparison to TDX. First, we validated both
 102 elevation models by calculating the height difference by subtracting the GPS elevation from the
 103 corresponding DEM pixel ($\Delta h = h_{DEM} - h_{GPS}$). Here, we understood and assessed accuracy in the
 104 same way as Wessel et al. (2018), where the systematic error is estimated by a statistical bias and
 105 the random error is estimated by the deviation in the height difference. We excluded height
 106 differences greater than 3 deviations before calculating the statistics. Considering a normal
 107 distribution, we assess the error by calculating the

108 mean error (ME),
$$ME = \frac{1}{n} \sum_{i=1}^n \Delta h_i;$$

109 root mean square error (RMSE),
$$RMSE = \sqrt{\frac{1}{n} \sum_{i=1}^n \Delta h_i^2};$$

110 and standard deviation (STD),
$$STD = \sqrt{\frac{1}{n-1} \sum_{i=1}^n (\Delta h_i - ME)^2}.$$

111 Other measures for accuracy assessment with non-normal error distributions were also calculated
 112 as proposed by Höhle and Höhle (2009), where $m_{\Delta h}$ is the median, i.e., 50% quantile:

113 median absolute deviation (MAD),
$$MAD = median_j(|\Delta h_j - m_{\Delta h}|);$$

114 the normalized median (NMAD),
$$NAMD = 1.4826 \cdot median_j(|\Delta h_j - m_{\Delta h}|);$$

115 and the absolute deviation at the 90% quantile (LE90) or linear error at the 90th percentile
 116 confidence level,
$$LE90 = \dot{Q}_{|\Delta h|}(0.9).$$

117 *REMA data.* REMA was constructed from stereoscopic imagery collected by four
 118 commercial satellites operated by DigitalGlobe Inc., with submeter resolution. The high spatial
 119 and radiometric resolutions of these imagers enable high-quality elevation extraction over low-
 120 contrast surfaces, such as snow cover and ice sheet interiors/accumulation zones. The REMA
 121 mosaic presents 68th and 90th percentile errors of 0.63 and 1.00 meters, respectively (Howat et
 122 al., 2019). The acquired strip DEMs that composed the tiles were collected between 2009 and
 123 2017; however, most of the tiles were collected in 2015 and 2016. The tiles are delivered in polar
 124 stereographic projection with a posting resolution of 8 meters. We used the REMA grid as a
 125 common grid reference for all reprojection and processing steps.

126 *TanDEM-X data.* The German Aerospace Center delivers TDX tiles in the geographic
 127 coordinates projection, posting in degrees resolution equivalent to an approximately 12x6 m
 128 resolution. First, we mosaicked the tiles to cover the Union Glacier (UG) basin and reprojected it
 129 to a polar stereographic projection (EPSG:3031) with bilinear resampling to REMA's common
 130 grid with an 8 m resolution. TDX images from Antarctica were acquired during the austral
 131 winter, between May and July 2013, and during the same months in 2014 (Rizzoli, Martone,
 132 Gonzalez, et al., 2017). The global validation performed by these authors confirmed an absolute

133 height accuracy of 3.49 m at the 90% confidence level, which is well below the 10 m mission
134 specification. A second assessment with GPS data with an accuracy of less than 0.5 m confirmed
135 an even lower error of below 2 m (Wessel et al., 2018). Because these accuracy assessments did
136 not include Antarctica, we assessed the height error annotated within the height error map (HEM)
137 of the TDX product, as the HEM is a good estimate of the theoretical random height error
138 (Wessel et al., 2018).

139 *DEM differences.* The time interval between both DEMs is between 2 and 4 years. Rivera
140 et al. (2014) reported a mean local elevation change at the narrow flux gate of -0.012 m a^{-1} , an
141 amount close to the estimated error of the measurements, which also indicates near-equilibrium
142 conditions. We also confirmed no elevation changes between 2005 and 2009 from the ICESat-
143 GLAS passes, with differences of less than one meter. For this reason, we performed a
144 comparison between both DEMs considering no surface elevation change. We subtracted REMA
145 from TDX. Because of the penetration of the SAR signal in the snow and the influence of surface
146 roughness, high-frequency noise is inherited in the TDX product. Each REMA strip DEM that is
147 composited with the mosaic product is filtered and downsampled to a 32 m grid during the
148 CryoSat-2 registration process (Howat et al., 2019). Therefore, we filtered the TDX using a
149 moving average filter with a window of 5x5 pixels (40x40 m), and we refer to this as TDX_A5.
150 The result is a smoother surface similar to that of REMA, and the different height map presented
151 in the results section is more homogeneous, representing a regional tendency that is higher or
152 lower than REMA.

153 *Masked areas.* As we aimed to assess higher accumulation rate zones, where the SAR
154 signal presented higher penetration, we masked the well-delimited bare ice surfaces with the blue
155 ice area (BIA) mask (Hui et al., 2014). We also mask areas with moderate to high slopes that are
156 greater than 5° ($\sim 10\%$). The mask polygons are depicted in Figure 1. The slope mask separated
157 the mountainous and rocky areas and simultaneously reduced the geometry's influence on the
158 SAR signal. BIAs in slope areas that were greater than 5° were considered to be steep areas.

159 *GPR data.* GPR profiles were collected using GSSI SIR[®] System-3000 with a 400 MHz
160 antenna in the 150 ns range (~ 12 m depth in dry snow) for the same 70 km displacement over UG
161 (Figure 1). We also collected data in the 600 ns range in the central valley to track the firn depth
162 up to 60 m. All data were collected in time mode and were not automatically synchronized with
163 GPS. The GPS acquisition was postprocessed in GIS software to precisely georeference and
164 topographically correct each GPR profile. We analyzed profiles where the firn/ice interface was
165 detectable and extracted the depth values of these transects.

166

167 3 Results

168 To validate and assess the REMA and TDX height accuracies, we compared the elevation
169 data with the GPS dataset. We used approximately 13 k grid points, and of these grid points,
170 ~10.5 k were on a flat snow surface, ~1.5 k were in BIAs exclusively, and ~1 k were in steep
171 areas. The statistics of the height differences are summarized in Table 1. In the supporting
172 information (SI), we presented the surface elevation profiles of the 3 datasets. The GPS ME
173 indicates an overall REMA offset of 1.01 with consistency through all classes. The profiles show
174 a smoother surface due to the rough terrain filter in the strip DEM registration (Howat et al.,
175 2019), and an excellent representation was confirmed, which reflected the low standard deviation.
176 The REMA has an STD (0.60 m) that was nearly half the RMSE (1.17 m). The STD is slightly
177 greater than the NAMD (0.45 m), indicating a good approximation to a normal distribution. The
178 NMAD is a more robust measure for the 68% probability level than the RMSE or standard
179 deviation. At the 90% probability level, the linear error LE90 is 1.73 m, which is greater than the
180 1.00 m reported by (Howat et al., 2019), but this result is expected for rougher terrain. Therefore,
181 we assumed that REMA is a good reference surface for further comparison to TDX.

182 The TDX presents a higher RMSE for all GPS points (3.27 m) and a great deviation in the
183 height differences (2.54 m). The NMAD and LE90 were 2.63 and 5.64 m, respectively. These
184 values are greater than the global absolute height accuracy of 3.49 m (Rizzoli, Martone,
185 Gonzalez, et al., 2017), which is clearly influenced by BIA and shallow snow/ice layer. There is a
186 divergence between each considered class due to a change in backscattering in each class. The
187 ME error of 5.04 m and low NMAD in BIAs indicates a consistent offset of the elevation over
188 these areas, where the calculated height values are offset by the constant correctness in the
189 registration step of the mosaic. The TDX profiles presented high-frequency variances in
190 elevation. These variances are observed in some of the GPS profiles with smaller amplitudes,
191 e.g., ‘Long-Driscoll’ profile (Figure S25). This high frequency represents the surface roughness,
192 but in TDX, the high frequency is enhanced by subsurface layering structures. The filtered DEM
193 (TDX_A5) presented a small decrease in all statistical parameters but maintained a close standard
194 deviation compared to the nonfiltered DEM.

195 The TDX profiles have a consistent positive difference over the BIAs. The same is
196 observed in a surrounding buffer zone, which is probably due to the shallowness of the glacial
197 ice. This bias is approximately 5 m and reflects the positive offset correction applied to the mean
198 phase center of the backscattered signal. Otherwise, the TDX shows agreement with the GPS
199 profile rather than with REMA, with the exception being in zones where the TDX gets deeper
200 than the GPS values, which we proposed to be related to the high accumulation zones.

202 **Table 1. (GPS POINTS Table)** Height accuracy assessments of each DEM against the field
 203 GPS. TDX_A5 is the average 5x5 filtered DEM. **(TDX-REMA Table)** Model-based
 204 accuracy analysis of TDX with REMA. Notably, we considered the TDX_A5 to be
 205 equivalent to a smoother REMA. A statistical summary was calculated in the classes of flat
 206 snow, BIA and steep areas as described in the text.

GPS POINTS							
Class	Points	ME (m)	RMS E (m)	STD (m)	MAD (m)	NMA D (m)	LE90 (m)
flat snow	10535	0.99	1.12	0.53	0.30	0.45	1.66
BIA	1511	1.05	1.18	0.55	0.23	0.35	1.59
steep	886	1.03	1.53	1.12	0.58	0.86	2.50
REMA - ALL	13326	1.01	1.17	0.60	0.31	0.45	1.73
flat snow	10560	1.55	2.75	2.27	1.46	2.16	4.97
BIA	1509	5.18	5.34	1.29	0.67	0.99	6.63
steep	1025	2.52	4.24	3.41	1.98	2.93	7.55
TDX - ALL	13503	2.05	3.27	2.54	1.78	2.63	5.64
flat snow	10562	1.29	2.51	2.15	1.04	1.55	4.81
BIA	1511	5.20	5.33	1.16	0.49	0.73	6.49
steep	998	2.61	4.27	3.38	1.17	1.74	8.11
TDX_A5 - ALL	13480	1.82	3.09	2.49	1.33	1.98	5.49
TDX_A5 - REMA							
Class	Points	ME (m)	RMS E (m)	STD (m)	MAD (m)	NMA D (m)	LE90 (m)
flat snow	107360150	-0.07	1.29	1.29	0.79	1.17	1.53
BIA	4084824	4.64	4.92	1.66	0.98	1.45	6.40
steep	12815811	2.93	5.12	4.20	2.86	4.24	8.57
ALL UG pixels	123434747	0.37	2.12	2.09	0.92	1.37	2.83

207

208 The DEM validation showed that REMA can be used as a reference surface. It is possible
 209 to compare TDX and REMA and assume that the spatial difference between both is a function of
 210 the X-band attenuation depth of the TDX data. This model-based accuracy assessment is
 211 summarized in (TDX_A5-REMA) Table 1, and the histogram of the height differences is
 212 presented in the SI (Figure S6). All results were calculated with the original TDX, not the filtered
 213 TDX. For all pixels, the histograms approximate a normal distribution, which is narrower but
 214 tailed to positive values for the contributions of mountain areas and BIAs. The mean height
 215 difference of 0.37 m indicates that TDX is close to REMA, and considering only flat snow cover
 216 areas, the mean difference is only -0.07 m.